

Dynamic Model of Curved Rods

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Abstract. The dynamic model of curved rod is obtained as an approximation of oscillations of three-dimensional curved rod-like linearized elastic body. The corresponding convergence result is stated.

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1. Introduction

In this paper we derive the evolution model of curved rod. The corresponding convergence result is stated. A similar problem in case of plates has been considered by Raoult [6], while Xiao Li-ming [5] derived and justified the evolution model of shells. The method used is a version of the asymptotic expansion method already applied to derivation of equilibrium models of curved rod [2, 3]. Problems and results are stated precisely but the proofs are mainly omitted; details can be found in [8].

By $(\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3)$ we denote the canonical basis in \mathbb{R}^3 . Vectors, vector-valued functions, matrices and matrix-valued functions are denoted by boldface letters. Euclidean inner product and norm of vectors in \mathbb{R}^n is denoted by \cdot and $|\cdot|$, respectively. Repeated index convention is accepted. The dual space of a Hilbert space H is denoted by H' . The same symbol $'$ stands for the derivative with respect to the longitudinal variable; partial derivatives with respect to space and time variables are denoted by ∂_i , $i = 1, 2, 3$ and ∂_t , ∂_{tt} , respectively.

2. The curved rod

Let \mathcal{C} be a simple regular curve in \mathbb{R}^3 defined by its natural parametrization $\Phi \in C^3([0, \ell]; \mathbb{R}^3)$. The tangent on \mathcal{C} at $\Phi(y^1)$ is defined by $\mathbf{t}(y^1) = \Phi'(y^1)$; obviously $\mathbf{t} \in C^2([0, \ell]; \mathbb{R}^3)$. In [4] it is proved that there exists a matrix-valued function $\mathbf{Q} \in C^2([0, \ell]; \text{SO}(3))$ such that

$$\mathbf{Q}\mathbf{e}_1 = \mathbf{t} \quad \text{on } [0, \ell].$$

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Then the local basis $\{\mathbf{t}(y^1), \mathbf{n}(y^1), \mathbf{b}(y^1)\}$ on \mathcal{C} at $\Phi(y^1)$ can be defined by

$$\mathbf{t}(y^1) = \mathbf{Q}(y^1)\mathbf{e}_1, \quad \mathbf{n}(y^1) = \mathbf{Q}(y^1)\mathbf{e}_2, \quad \mathbf{b}(y^1) = \mathbf{Q}(y^1)\mathbf{e}_3, \quad y^1 \in [0, \ell].$$

Smoothness of \mathbf{Q} implies that $\mathbf{t}, \mathbf{n}, \mathbf{b} \in C^2([0, \ell]; \mathbb{R}^3)$. Note that this basis exists even though the smooth Frenet frame does not exist globally. Let

$$\mathbf{R}(y^1) = \left(\frac{d}{dy^1} \mathbf{Q}^T(y^1) \right) \mathbf{Q}(y^1), \quad y^1 \in [0, \ell].$$

Matrix $\mathbf{R}(y^1)$ is antisymmetric for all $y^1 \in [0, \ell]$. Moreover, the following generalization of the Frenet equations holds:

$$\frac{d}{dy^1} \mathbf{Q}^T = \mathbf{R} \mathbf{Q}^T.$$

In case of generic curve ($|\Phi''| > 0$), matrix-valued function \mathbf{Q} can be chosen so that $(\mathbf{t}, \mathbf{n}, \mathbf{b})$ and the Frenet frame coincide. In this case it holds

$$\mathbf{R} = \begin{pmatrix} 0 & \kappa & 0 \\ -\kappa & 0 & \tau \\ 0 & -\tau & 0 \end{pmatrix},$$

where κ is the curvature and τ is the torsion of the curve \mathcal{C} .

Now we define the domain in \mathbb{R}^3 which represents the curved rod. Let $\varepsilon > 0$ be a small parameter and let $S \subset \mathbb{R}^2$ be a bounded domain with Lipschitz boundary; points in S are denoted by (z^2, z^3) . The coordinates z^2 and z^3 can be chosen so that $(0, 0)$ is the center of mass of S , i.e.,

$$\int_S z^2 dz^2 dz^3 = \int_S z^3 dz^2 dz^3 = 0. \quad (1)$$

Let

$$\Omega^\varepsilon = (0, \ell) \times \varepsilon S, \quad \Gamma^\varepsilon = (0, \ell) \times \varepsilon \partial S, \quad B_{y^1}^\varepsilon = \{y^1\} \times \varepsilon S.$$

Γ^ε is the lateral boundary, ℓ is the length, $B_{y^1}^\varepsilon$ is the cross-section at $y^1 \in [0, \ell]$, while ε is the thickness of the cylinder Ω^ε . B_0^ε and B_ℓ^ε are called the bases of the rod. Let $\mathbf{P} : \overline{\Omega}^\varepsilon \rightarrow \mathbb{R}^3$ be defined by

$$\mathbf{P}(y) = \Phi(y^1) + y^2 \mathbf{n}(y^1) + y^3 \mathbf{b}(y^1), \quad y^1 \in [0, \ell], \quad (y^2, y^3) \in \overline{\varepsilon S}.$$

Since \mathcal{C} is simple and regular curve, there exists $\varepsilon_0 > 0$ such that the function \mathbf{P} is injective; especially one has

$$g(y) = (\det \nabla \mathbf{P}(y))^2 = (1 - R_{12}(y^1)y^2 - R_{13}(y^1)y^3)^2 > 0, \quad y \in \Omega^\varepsilon.$$

In the sequel we restrict ourselves to $\varepsilon \in (0, \varepsilon_0)$. The domain, see Figure 1,

$$\tilde{\Omega}^\varepsilon = \mathbf{P}(\Omega^\varepsilon)$$

is called the curved rod; its lateral boundary and its cross-sections at $\Phi(y^1)$ are, respectively,

$$\tilde{\Gamma}^\varepsilon = \mathbf{P}(\Gamma^\varepsilon), \quad \tilde{B}_{y^1}^\varepsilon = \mathbf{P}(B_{y^1}^\varepsilon).$$

\tilde{B}_0^ε and $\tilde{B}_\ell^\varepsilon$ are the bases of $\tilde{\Omega}^\varepsilon$. Because of (1), the curve \mathcal{C} passes through the centers of mass of $\tilde{B}_{y^1}^\varepsilon$, $y^1 \in [0, \ell]$, thus we say that \mathcal{C} is the middle curve of the curved rod $\tilde{\Omega}^\varepsilon$.

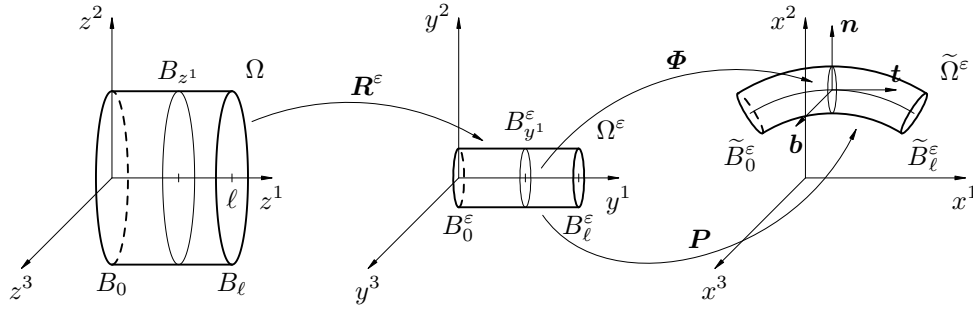


Figure 1. Canonical domain, thin cylinder and curved rod.

3. The three-dimensional evolution equations

We assume that $\tilde{\Omega}^\varepsilon$ is the natural state of a linearized isotropic elastic body with the Lamé constants λ and μ . Let $\tilde{\Omega}^\varepsilon$ be clamped on its bases \tilde{B}_0^ε and $\tilde{B}_\ell^\varepsilon$ and force-free on its lateral boundary $\tilde{\Gamma}^\varepsilon$, and let $\tilde{\mathbf{U}}_0^\varepsilon$ and $\tilde{\mathbf{U}}_1^\varepsilon$ be the initial displacement and velocity of $\tilde{\Omega}^\varepsilon$, respectively. Let $T > 0$. Then the oscillations of $\tilde{\Omega}^\varepsilon$ are described by a function $\tilde{\mathbf{U}}^\varepsilon$ which is formally a solution of

$$\begin{aligned} \tilde{\varrho}^\varepsilon \partial_{tt} \tilde{\mathbf{U}}^\varepsilon - \operatorname{div}(\lambda \operatorname{tr} \mathbf{e}(\tilde{\mathbf{U}}^\varepsilon) \mathbf{I} + 2\mu \mathbf{e}(\tilde{\mathbf{U}}^\varepsilon)) &= \tilde{\mathbf{F}}^\varepsilon \quad \text{in } (0, T) \times \tilde{\Omega}^\varepsilon, \\ \tilde{\mathbf{U}}^\varepsilon &= 0 \quad \text{on } (0, T) \times (\tilde{B}_0^\varepsilon \cup \tilde{B}_\ell^\varepsilon), \\ \lambda \operatorname{tr}(\mathbf{e}(\tilde{\mathbf{U}}^\varepsilon)) \tilde{\boldsymbol{\nu}} + 2\mu \mathbf{e}(\tilde{\mathbf{U}}^\varepsilon) \tilde{\boldsymbol{\nu}} &= 0 \quad \text{on } (0, T) \times \tilde{\Gamma}^\varepsilon, \\ \tilde{\mathbf{U}}^\varepsilon|_{t=0} &= \tilde{\mathbf{U}}_0^\varepsilon, \quad \partial_t \tilde{\mathbf{U}}^\varepsilon|_{t=0} = \tilde{\mathbf{U}}_1^\varepsilon \quad \text{in } \tilde{\Omega}^\varepsilon. \end{aligned}$$

Here $\tilde{\mathbf{F}}^\varepsilon$ is the volume force density acting on the curved rod, $\tilde{\varrho}^\varepsilon$ is the mass density of the rod, $\tilde{\boldsymbol{\nu}}$ is the unit outer normal at the boundary of the curved rod, and $\mathbf{e}(\tilde{\mathbf{V}})$ denotes the symmetrized gradient of the function $\tilde{\mathbf{V}}$, i.e.,

$$\mathbf{e}(\tilde{\mathbf{V}}) = \frac{1}{2}(\nabla \tilde{\mathbf{V}} + (\nabla \tilde{\mathbf{V}})^T).$$

More precisely, let us introduce the function space

$$\mathcal{V}(\tilde{\Omega}^\varepsilon) = \{\tilde{\mathbf{V}} \in H^1(\tilde{\Omega}^\varepsilon)^3 \mid \tilde{\mathbf{V}}|_{\tilde{B}_0^\varepsilon} = \tilde{\mathbf{V}}|_{\tilde{B}_\ell^\varepsilon} = 0\},$$

which is a Hilbert space for the scalar product of $H^1(\Omega)^3$. Differential equations are then formally equivalent to:

$$\begin{aligned} \int_{\tilde{\Omega}^\varepsilon} \tilde{\varrho}^\varepsilon \partial_{tt} \tilde{\mathbf{U}}^\varepsilon \cdot \tilde{\mathbf{V}} \, dx + \int_{\tilde{\Omega}^\varepsilon} \lambda \operatorname{tr} \mathbf{e}(\tilde{\mathbf{U}}^\varepsilon) \operatorname{tr} \mathbf{e}(\tilde{\mathbf{V}}) + 2\mu \mathbf{e}(\tilde{\mathbf{U}}^\varepsilon) \cdot \mathbf{e}(\tilde{\mathbf{V}}) \, dx \\ = \int_{\tilde{\Omega}^\varepsilon} \tilde{\mathbf{F}}^\varepsilon \cdot \tilde{\mathbf{V}} \, dx, \quad \tilde{\mathbf{V}} \in \mathcal{V}(\tilde{\Omega}^\varepsilon), \quad 0 < t < T, \end{aligned} \quad (2)$$

$$\tilde{\mathbf{U}}^\varepsilon|_{t=0} = \tilde{\mathbf{U}}_0^\varepsilon, \quad \partial_t \tilde{\mathbf{U}}^\varepsilon|_{t=0} = \tilde{\mathbf{U}}_1^\varepsilon. \quad (3)$$

Lemma 1. *Let $\tilde{\mathbf{F}}^\varepsilon \in L^2(0, T; L^2(\tilde{\Omega}^\varepsilon)^3)$, $\tilde{\mathbf{U}}_0^\varepsilon \in \mathcal{V}(\tilde{\Omega}^\varepsilon)$ and $\tilde{\mathbf{U}}_1^\varepsilon \in L^2(\tilde{\Omega}^\varepsilon)^3$. Then there exists a unique solution $\tilde{\mathbf{U}}^\varepsilon$ of the problem (2), (3), such that*

$$\tilde{\mathbf{U}}^\varepsilon \in C([0, T]; \mathcal{V}(\tilde{\Omega}^\varepsilon)), \quad \partial_t \tilde{\mathbf{U}}^\varepsilon \in C([0, T]; L^2(\tilde{\Omega}^\varepsilon)), \quad \partial_{tt} \tilde{\mathbf{U}}^\varepsilon \in L^2(0, T; \mathcal{V}(\tilde{\Omega}^\varepsilon)').$$

At the end of this section we rewrite (2), (3) in curvilinear coordinates defined by \mathbf{P} . Covariant basis of the curved rod is defined by

$$\mathbf{g}_i = \partial_i \mathbf{P} : \Omega^\varepsilon \rightarrow \mathbb{R}^3, \quad i = 1, 2, 3.$$

Vectors $(\mathbf{g}^1, \mathbf{g}^2, \mathbf{g}^3)$ satisfying

$$\mathbf{g}^j \cdot \mathbf{g}_i = \delta_i^j \quad \text{on } \tilde{\Omega}^\varepsilon, \quad i, j = 1, 2, 3,$$

where δ_i^j is the Kronecker symbol, form the contravariant basis on $\tilde{\Omega}^\varepsilon$. The contravariant metric tensor $\mathbf{G} = (g^{ij})$ and the Christoffel symbols Γ_{jk}^i of the curved rod $\tilde{\Omega}^\varepsilon$ are defined by

$$g^{ij} = \mathbf{g}^i \cdot \mathbf{g}^j, \quad \Gamma_{jk}^i = \mathbf{g}^i \cdot \partial_j \mathbf{g}_k \quad \text{on } \tilde{\Omega}^\varepsilon, \quad i, j, k = 1, 2, 3.$$

The corresponding function space to $\mathcal{V}(\tilde{\Omega}^\varepsilon)$ is the space

$$\mathcal{V}(\Omega^\varepsilon) = \{\mathbf{V} \in H^1(\Omega^\varepsilon)^3 \mid \mathbf{V}|_{B_0^\varepsilon} = \mathbf{V}|_{B_\ell^\varepsilon} = 0\}.$$

The displacements and velocities are rewritten in contravariant basis, while the force density in covariant basis

$$\begin{aligned} \tilde{\mathbf{U}}^\varepsilon \circ \mathbf{P} = U_i^\varepsilon \mathbf{g}^i, \quad \tilde{\mathbf{U}}_0^\varepsilon \circ \mathbf{P} = (U_0^\varepsilon)_i \mathbf{g}^i, \quad \tilde{\mathbf{U}}_1^\varepsilon \circ \mathbf{P} = (U_1^\varepsilon)_i \mathbf{g}^i, \\ \tilde{\mathbf{V}} \circ \mathbf{P} = V_i \mathbf{g}^i, \quad \tilde{\mathbf{F}}^\varepsilon \circ \mathbf{P} = F^{\varepsilon i} \mathbf{g}_i. \end{aligned}$$

We define vector functions

$$\mathbf{U}^\varepsilon = U_i^\varepsilon \mathbf{e}_i, \quad \mathbf{U}_0^\varepsilon = (U_0^\varepsilon)_i \mathbf{e}_i, \quad \mathbf{U}_1^\varepsilon = (U_1^\varepsilon)_i \mathbf{e}_i, \quad \mathbf{V} = V_i \mathbf{e}_i, \quad \mathbf{F}^\varepsilon = F^{\varepsilon i} \mathbf{e}_i.$$

Let $\varrho^\varepsilon = \tilde{\varrho}^\varepsilon \circ \mathbf{P}$ and

$$\begin{aligned}\boldsymbol{\gamma}(\mathbf{V}) &= \mathbf{e}(\mathbf{V}) - V_i \boldsymbol{\Gamma}^i, \\ \mathbf{A}\boldsymbol{\sigma}^1 \cdot \boldsymbol{\sigma}^2 &= \lambda \operatorname{tr} \boldsymbol{\sigma}^1 \operatorname{tr} \boldsymbol{\sigma}^2 + 2\mu \boldsymbol{\sigma}^1 \cdot \boldsymbol{\sigma}^2, \quad \boldsymbol{\sigma}^1, \boldsymbol{\sigma}^2 \in \operatorname{Sym}(\mathbb{R}^3),\end{aligned}$$

where $\operatorname{Sym}(\mathbb{R}^3)$ denotes the space of all real symmetric matrices of order 3. The system (2), (3) is then equivalent to the following equation of motion in curvilinear coordinates:

$$\begin{aligned}\int_{\Omega^\varepsilon} \varrho^\varepsilon \mathbf{G} \partial_{tt} \mathbf{U}^\varepsilon \cdot \mathbf{V} \sqrt{g} dy + \int_{\Omega^\varepsilon} \mathbf{A} \boldsymbol{\gamma}(\mathbf{U}^\varepsilon) \cdot \boldsymbol{\gamma}(\mathbf{V}) \sqrt{g} dy \\ = \int_{\Omega^\varepsilon} \mathbf{F}^\varepsilon \cdot \mathbf{V} \sqrt{g} dy, \quad \mathbf{V} \in \mathcal{V}(\Omega^\varepsilon), \quad 0 < t < T,\end{aligned}\tag{4}$$

$$\mathbf{U}^\varepsilon|_{t=0} = \mathbf{U}_0^\varepsilon, \quad \partial_t \mathbf{U}^\varepsilon|_{t=0} = \mathbf{U}_1^\varepsilon.\tag{5}$$

4. A priori estimates on fixed domain

Our main goal is to find the limit of $(\tilde{\mathbf{U}}^\varepsilon, \varepsilon > 0)$ as ε tends to zero, as well as the equations satisfied by the limit. Problems for both $\tilde{\mathbf{U}}^\varepsilon$ and \mathbf{U}^ε are posed on ε -dependent domains. Now we transform the problem (4), (5) to ε -independent domain, see Figure 1. As a consequence, the coefficients of the resulting weak formulation will depend on ε explicitly, and calculation of the limit will be enabled.

Let $\Omega = (0, \ell) \times S$, and let $\mathbf{R}^\varepsilon : \bar{\Omega} \rightarrow \mathbb{R}^3$ be defined by

$$\mathbf{R}^\varepsilon(z) = (z^1, \varepsilon z^2, \varepsilon z^3), \quad z \in \Omega, \quad \varepsilon \in (0, \varepsilon_0).$$

By Γ and S_{z^1} we denote the lateral surface of Ω and its cross-section at $z^1 \in [0, \ell]$, respectively. To the functions $\mathbf{U}^\varepsilon, \mathbf{U}_0^\varepsilon, \mathbf{U}_1^\varepsilon, \mathbf{F}^\varepsilon, g, \mathbf{g}_i, \mathbf{g}^i, \varrho^\varepsilon, \mathbf{G}, \mathbf{R}, \Gamma_{jk}^i, i, j, k = 1, 2, 3$, defined on Ω^ε , we associate the functions $\mathbf{u}(\varepsilon), \mathbf{u}_0(\varepsilon), \mathbf{u}_1(\varepsilon), \mathbf{f}(\varepsilon), g(\varepsilon), \mathbf{g}_i(\varepsilon), \mathbf{g}^i(\varepsilon), \varrho(\varepsilon), \mathbf{G}(\varepsilon), \mathbf{R}(\varepsilon), \Gamma_{jk}^i(\varepsilon), i, j, k = 1, 2, 3$, defined on Ω by composition with \mathbf{R}^ε . Let

$$\mathcal{V}(\Omega) = \left\{ \mathbf{v} = (v_1, v_2, v_3) \in H^1(\Omega)^3 \mid \mathbf{v}|_{B_0} = \mathbf{v}|_{B_\ell} = \mathbf{0} \right\},$$

$$a(\varepsilon) : \mathcal{V}(\Omega) \times \mathcal{V}(\Omega) \rightarrow \mathbb{R}, \quad a(\varepsilon)(\mathbf{v}, \mathbf{w}) = \int_{\Omega} \mathbf{A}(\varepsilon) \frac{1}{\varepsilon} \boldsymbol{\gamma}^\varepsilon(\mathbf{v}) \cdot \frac{1}{\varepsilon} \boldsymbol{\gamma}^\varepsilon(\mathbf{w}) \sqrt{g(\varepsilon)} dz,$$

$$b(\varepsilon) : L^2(\Omega)^3 \times L^2(\Omega)^3 \rightarrow \mathbb{R}, \quad b(\varepsilon)(\mathbf{v}, \mathbf{w}) = \int_{\Omega} \frac{1}{\varepsilon^2} \varrho(\varepsilon) \mathbf{G}(\varepsilon) \mathbf{v} \cdot \mathbf{w} \sqrt{g(\varepsilon)} dz,$$

$$\boldsymbol{\gamma}^\varepsilon(\mathbf{v}) = \frac{1}{\varepsilon} \boldsymbol{\gamma}_z(\mathbf{v}) + \boldsymbol{\gamma}_y(\mathbf{v}) - v_i \boldsymbol{\Gamma}^i(\varepsilon),$$

$$\boldsymbol{\gamma}_y(\mathbf{v}) = \begin{pmatrix} \partial_1 v_1 & \frac{1}{2} \partial_1 v_2 & \frac{1}{2} \partial_1 v_3 \\ \frac{1}{2} \partial_1 v_2 & 0 & 0 \\ \frac{1}{2} \partial_1 v_3 & 0 & 0 \end{pmatrix},$$

$$\boldsymbol{\gamma}_z(\mathbf{v}) = \begin{pmatrix} 0 & \frac{1}{2}\partial_2 v_1 & \frac{1}{2}\partial_3 v_1 \\ \frac{1}{2}\partial_2 v_1 & \partial_2 v_2 & \frac{1}{2}(\partial_2 v_3 + \partial_3 v_2) \\ \frac{1}{2}\partial_3 v_1 & \frac{1}{2}(\partial_2 v_3 + \partial_3 v_2) & \partial_3 v_3 \end{pmatrix},$$

$$\mathbf{A}(\varepsilon)\boldsymbol{\sigma}^1 \cdot \boldsymbol{\sigma}^2 = \lambda \operatorname{tr} \mathbf{G}(\varepsilon)\boldsymbol{\sigma}^1 \operatorname{tr} \mathbf{G}(\varepsilon)\boldsymbol{\sigma}^2 + 2\mu \mathbf{G}(\varepsilon)\boldsymbol{\sigma}^1 \mathbf{G}(\varepsilon) \cdot \boldsymbol{\sigma}^2, \quad \boldsymbol{\sigma}^1, \boldsymbol{\sigma}^2 \in \operatorname{Sym}(\mathbb{R}^3).$$

The problem (4), (5) is then equivalent to the following problem:

Find $\mathbf{u}(\varepsilon)$ such that

$$\frac{d}{dt}b(\varepsilon)(\partial_t \mathbf{u}(\varepsilon), \mathbf{v}) + a(\varepsilon)(\mathbf{u}(\varepsilon), \mathbf{v}) = \left(\frac{1}{\varepsilon^2} \mathbf{f}(\varepsilon) \sqrt{g(\varepsilon)} \mid \mathbf{v} \right)_{L^2(\Omega)^3},$$

$$\mathbf{v} \in \mathcal{V}(\Omega), \quad 0 < t < T, \quad (6)$$

$$\mathbf{u}(\varepsilon)|_{t=0} = \mathbf{u}_0(\varepsilon), \quad \partial_t \mathbf{u}(\varepsilon)|_{t=0} = \mathbf{u}_1(\varepsilon). \quad (7)$$

The family $(\mathbf{u}(\varepsilon), 0 < \varepsilon < \varepsilon_0)$ satisfies

$$\mathbf{u}(\varepsilon) \in C([0, T]; \mathcal{V}(\Omega)), \quad \partial_t \mathbf{u}(\varepsilon) \in C([0, T]; L^2(\Omega)^3), \quad \partial_{tt} \mathbf{u}(\varepsilon) \in L^2(0, T; \mathcal{V}(\Omega)').$$

In general, this family is not bounded with respect to ε . The conditions ensuring certain uniform boundedness are stated in the following theorem.

Theorem 1 (a priori estimates). *Let us assume that*

$$\mathbf{f}(\varepsilon) = \varepsilon^2 \mathbf{f}, \quad \varrho(\varepsilon) = \varepsilon^2 \varrho, \quad (8)$$

$$\left\| \frac{1}{\varepsilon} \boldsymbol{\gamma}^\varepsilon(\mathbf{u}_0(\varepsilon)) \right\|_{L^2(\Omega)^9} \leq C_u, \quad \mathbf{u}_1(\varepsilon) \rightharpoonup \mathbf{u}_1 \quad \text{weakly in } L^2(\Omega)^3, \quad (9)$$

where ϱ and C_u are constants independent of ε . Then there exists $C > 0$ such that

$$\|\mathbf{u}(\varepsilon)\|_{\mathcal{V}(\Omega)} \leq C, \quad \|\partial_t \mathbf{u}(\varepsilon)\|_{L^2(\Omega)^3} \leq C, \quad \left\| \frac{1}{\varepsilon} \boldsymbol{\gamma}^\varepsilon(\mathbf{u}(\varepsilon)) \right\|_{L^2(\Omega)^9} \leq C, \quad \varepsilon \in (0, \varepsilon_0). \quad (10)$$

Proof of Theorem 1 is long and we omit it here; details can be found in [8]. A simple consequence of a priori estimate (10) and smoothness of $\mathbf{u}(\varepsilon)$ is the following convergence result.

Corollary 1. *There is a subsequence of $(\mathbf{u}(\varepsilon), \varepsilon > 0)$ (still denoted by ε) and functions \mathbf{u} and $\boldsymbol{\gamma}$, $\mathbf{u} \in L^\infty(0, T; \mathcal{V}(\Omega))$, $\partial_t \mathbf{u} \in L^\infty(0, T; L^2(\Omega)^3)$, $\boldsymbol{\gamma} \in L^\infty(0, T; L^2(\Omega)^9)$, such that*

$$\begin{aligned} \mathbf{u}(\varepsilon) &\overset{*}{\rightharpoonup} \mathbf{u} \quad \text{weak * in } L^\infty(0, T; \mathcal{V}(\Omega)), \\ \partial_t \mathbf{u}(\varepsilon) &\overset{*}{\rightharpoonup} \partial_t \mathbf{u} \quad \text{weak * in } L^\infty(0, T; L^2(\Omega)^3), \\ \frac{1}{\varepsilon} \boldsymbol{\gamma}^\varepsilon(\mathbf{u}(\varepsilon)) &\overset{*}{\rightharpoonup} \boldsymbol{\gamma} \quad \text{weak * in } L^\infty(0, T; L^2(\Omega)^9). \end{aligned}$$

5. The limit problem

The relevant function spaces for description of the limit problem are the following Hilbert spaces

$$\begin{aligned} V_0 &= H_0^1(0, \ell) \times H_0^2(0, \ell) \times H_0^2(0, \ell), \quad \mathcal{V}_0 = \{\mathbf{v} \in V_0 \mid v_1' - R_{12}v_2 - R_{13}v_3 = 0\}, \\ H_0 &= L^2(0, \ell)^3, \quad \mathcal{H}_0 = \{\mathbf{v} \in H_0^1(0, \ell) \times L^2(0, \ell) \times L^2(0, \ell) \mid v_1' - R_{12}v_2 - R_{13}v_3 = 0\}, \\ W_0 &= V_0 \times H_0^1(0, \ell), \quad \mathcal{W}_0 = \mathcal{V}_0 \times H_0^1(0, \ell), \end{aligned}$$

with norms

$$\begin{aligned} \|\mathbf{v}\|_{V_0}^2 &= \|v_1'\|_{L^2(0, \ell)}^2 + \|v_2''\|_{L^2(0, \ell)}^2 + \|v_3''\|_{L^2(0, \ell)}^2, \\ \|\mathbf{v}\|_{H_0}^2 &= \|v_1\|_{L^2(0, \ell)}^2 + \|v_2\|_{L^2(0, \ell)}^2 + \|v_3\|_{L^2(0, \ell)}^2, \\ \|\mathbf{v}\|_{W_0}^2 &= \|\mathbf{v}\|_{V_0}^2 + \|\psi'\|_{L^2(0, \ell)}^2, \quad \mathbf{v} = (\mathbf{v}, \psi) \in W_0. \end{aligned}$$

V_0 is the usual function space for rods, but the limit displacement from Corollary 1 describes the inextensible oscillations of the rod, i.e., it belongs to the subspace \mathcal{V}_0 of V_0 . Also note that \mathcal{H}_0 is the closure of \mathcal{V}_0 in the norm of H_0 .

Let us introduce bilinear forms $a_0 : W_0 \times W_0 \rightarrow \mathbb{R}$ and $b_0 : H_0 \times H_0 \rightarrow \mathbb{R}$ by

$$a_0(\mathbf{v}, \mathbf{w}) := \int_0^\ell \mathbf{H} \mathbf{a}(\mathbf{v}) \cdot \mathbf{a}(\mathbf{w}) dz^1, \quad b_0(\mathbf{v}, \mathbf{w}) := (\varrho A \mathbf{v} \mid \mathbf{w})_{H_0},$$

where $\mathbf{a}(\mathbf{v})$ is defined for $\mathbf{v} = (\mathbf{v}, \psi) \in W_0$ by

$$\begin{aligned} a_1(\mathbf{v}) &= \psi' + R_{12}(v_3' + R_{13}v_1 + R_{23}v_2) - R_{13}(v_2' + R_{12}v_1 - R_{23}v_3), \\ a_2(\mathbf{v}) &= -(v_3' + R_{13}v_1 + R_{23}v_2)' + R_{12}\psi - R_{23}(v_2' + R_{12}v_1 - R_{23}v_3), \\ a_3(\mathbf{v}) &= -(v_2' + R_{12}v_1 - R_{23}v_3)' - R_{13}\psi + R_{23}(v_3' + R_{13}v_1 + R_{23}v_2). \end{aligned}$$

The positive definite symmetric matrix \mathbf{H} depends on geometry of the rod and on its elastic properties. Precisely, let the moments of inertia of the cross-section S and the area of S be denoted by

$$I_{23} = - \int_S z^2 z^3 dz^2 dz^3, \quad I_\alpha = \int_S (z^\alpha)^2 dz^2 dz^3, \quad \alpha = 2, 3, \quad A = \int_S dz^2 dz^3,$$

and let $p \in H^1(S)$ be the warping function, i.e., a unique solution of the problem

$$\Delta p = 0 \quad \text{in } S, \quad \frac{\partial p}{\partial \boldsymbol{\nu}} = \begin{pmatrix} z^3 \\ -z^2 \end{pmatrix} \cdot \boldsymbol{\nu} \quad \text{on } \partial S, \quad \int_S p dz^2 dz^3 = 0,$$

where $\boldsymbol{\nu}$ denotes the unit outer normal on S . It can be shown that the number

$$K = \int_S ((\partial_2 p - z^3)^2 + (\partial_3 p + z^2)^2) dz^2 dz^3$$

is positive; μK is called the torsion rigidity. Now,

$$\mathbf{H} = \begin{pmatrix} \mu K & 0 & 0 \\ 0 & EI_3 & -EI_{23} \\ 0 & -EI_{23} & EI_2 \end{pmatrix},$$

where E is the Young modulus.

Theorem 2. *Let us assume that*

$$\mathbf{u}_0(\varepsilon) \rightharpoonup \mathbf{u}_0 \quad \text{weakly in } \mathcal{V}(\Omega), \quad \mathbf{u}_1 \in \mathcal{H}_0. \quad (11)$$

Let \mathbf{u} and γ be the limits from Corollary 1. Then there exists $\phi \in L^\infty(0, T; H_0^1(0, \ell))$ such that $\mathbf{u} = (\mathbf{u}, \phi)$ satisfies

$$\begin{aligned} \mathbf{u} &\in L^\infty(0, T; \mathcal{V}_0), \quad \partial_t \mathbf{u} \in L^\infty(0, T; \mathcal{H}_0), \quad \partial_{tt} \mathbf{u} \in L^2(0, T; \mathcal{V}'_0), \\ \frac{d}{dt} b_0(\partial_t \mathbf{u}, \mathbf{v}) + a_0(\mathbf{u}, \mathbf{v}) &= (\bar{\mathbf{f}} | \mathbf{v})_{H_0}, \quad \mathbf{v} = (\mathbf{v}, \psi) \in \mathcal{W}_0, \end{aligned} \quad (12)$$

$$\mathbf{u}(0) = \mathbf{u}_0, \quad \partial_t \mathbf{u}(0) = \mathbf{u}_1, \quad (13)$$

where

$$\bar{\mathbf{f}} = \int_S \mathbf{f} dz^2 dz^3 \in L^2(0, T; \mathcal{H}_0).$$

Moreover, the limit function γ is of the form

$$\gamma = \begin{pmatrix} a_3(\mathbf{u})z^2 + a_2(\mathbf{u})z^3 & \cdot & \cdot \\ \frac{1}{2}a_1(\mathbf{u})(\partial_2 p - z^3) & -\nu(a_3(\mathbf{u})z^2 + a_2(\mathbf{u})z^3) & \cdot \\ \frac{1}{2}a_1(\mathbf{u})(\partial_3 p + z^2) & 0 & -\nu(a_3(\mathbf{u})z^2 + a_2(\mathbf{u})z^3) \end{pmatrix},$$

where ν is the Poisson ratio.

The problem (12), (13) is not a classical one, because there is no time derivative of ϕ . The equivalent form of (12) is

$$\begin{aligned} a_0((\mathbf{u}, \phi), (0, \psi)) &= 0, \quad \psi \in H_0^1(0, \ell), \\ \frac{d}{dt} b_0(\partial_t \mathbf{u}, \mathbf{v}) + a_0((\mathbf{u}, \phi), (\mathbf{v}, 0)) &= (\bar{\mathbf{f}} | \mathbf{v})_{H_0}, \quad \mathbf{v} \in \mathcal{V}_0. \end{aligned} \quad (14)$$

It follows that there exists a linear continuous operator $D : \mathcal{V}_0 \rightarrow H_0^1(0, \ell)$ such that

$$\phi = D\mathbf{u}. \quad (15)$$

From (14) it follows that \mathbf{u} is a solution of the following standard problem

$$\frac{d}{dt} b_0(\partial_t \mathbf{u}, \mathbf{v}) + d_0(\mathbf{u}, \mathbf{v}) = (\bar{\mathbf{f}} | \mathbf{v})_{H_0}, \quad \mathbf{v} \in \mathcal{V}_0, \quad (16)$$

$$\mathbf{u}(0) = \mathbf{u}_0, \quad \partial_t \mathbf{u}(0) = \mathbf{u}_1, \quad (17)$$

where

$$d_0 : \mathcal{V}_0 \times \mathcal{V}_0 \rightarrow \mathbb{R}, \quad d_0(\mathbf{v}, \mathbf{w}) = a_0((\mathbf{u}, 0), (\mathbf{v}, 0)) - a_0((0, D\mathbf{u}), (0, D\mathbf{v})).$$

Since bilinear forms b_0 and d_0 are elliptic, the classical theory of evolution equations implies that the problem (16), (17) has a unique solution \mathbf{u} such that

$$\mathbf{u} \in \mathcal{V}_0^c(0, T) = \{ \mathbf{v} \in C([0, T]; \mathcal{V}_0) \mid \partial_t \mathbf{v} \in C([0, T]; \mathcal{H}_0), \partial_{tt} \mathbf{v} \in L^2(0, T; \mathcal{V}'_0) \}.$$

The function ϕ is uniquely determined by (15) and $\phi \in C([0, T]; H_0^1(0, \ell))$. It follows that the limit functions \mathbf{u} and γ are unique, hence whole families $(\mathbf{u}(\varepsilon), 0 < \varepsilon < \varepsilon_0)$ and $((1/\varepsilon)\gamma^\varepsilon(\mathbf{u}(\varepsilon)), 0 < \varepsilon < \varepsilon_0)$ are convergent. Thus we proved

Theorem 3. *Let the function $\mathbf{u}(\varepsilon)$, for $0 < \varepsilon < \varepsilon_0$, be the solution of (6), (7), and let the assumptions (8), (9) and (11) be fulfilled. Then*

$$\mathbf{u}(\varepsilon) \in C([0, T]; \mathcal{V}(\Omega)), \quad \partial_t \mathbf{u}(\varepsilon) \in C([0, T]; L^2(\Omega)^3), \quad \partial_{tt} \mathbf{u}(\varepsilon) \in L^2(0, T; \mathcal{V}(\Omega)'),$$

and

$$\begin{aligned} \mathbf{u}(\varepsilon) &\overset{*}{\rightharpoonup} \mathbf{u} \quad \text{weak * in } L^\infty(0, T; \mathcal{V}(\Omega)), \\ \partial_t \mathbf{u}(\varepsilon) &\overset{*}{\rightharpoonup} \partial_t \mathbf{u} \quad \text{weak * in } L^\infty(0, T; L^2(\Omega)^3), \\ \partial_{tt} \mathbf{u}(\varepsilon) &\rightharpoonup \partial_{tt} \mathbf{u} \quad \text{weakly in } L^2(0, T; \mathcal{V}(\Omega)'), \\ \frac{1}{\varepsilon} \gamma^\varepsilon(\mathbf{u}(\varepsilon)) &\overset{*}{\rightharpoonup} \gamma \quad \text{weak * in } L^\infty(0, T; L^2(\Omega)^9), \end{aligned}$$

when $\varepsilon \rightarrow 0$, where $\mathbf{u} = (\mathbf{u}, \phi) \in \mathcal{V}_0^c(0, T) \times C([0, T]; H_0^1(0, \ell))$ is a unique solution of (12), (13) and γ as given in Theorem 2.

6. Dynamic curved rod model

The system (12), (13) is posed on the canonical domain Ω . The evolution equations of a curved rod in curvilinear coordinates follow from (12), (13), by the change of variables $\mathbf{R}^\varepsilon : \Omega \rightarrow \Omega^\varepsilon$ (see Figure 1). It is easy to show that

$$I_{23}^\varepsilon = \varepsilon^4 I_{23}, \quad I_\alpha^\varepsilon = \varepsilon^4 I_\alpha, \quad \alpha = 2, 3, \quad A^\varepsilon = \varepsilon^2 A,$$

are moments of inertia and area of εS , respectively. Also, the warping function p^ε of εS is given by $p^\varepsilon \circ \mathbf{R}^\varepsilon = \varepsilon^2 p$, so $K^\varepsilon = \varepsilon^4 K$. Thus the matrix

$$\mathbf{H}^\varepsilon = \begin{pmatrix} \mu K^\varepsilon & 0 & 0 \\ 0 & EI_3^\varepsilon & -EI_{23}^\varepsilon \\ 0 & -EI_{23}^\varepsilon & EI_2^\varepsilon \end{pmatrix} = \varepsilon^4 \mathbf{H}$$

is symmetric and positive definite. The forms $a_0^\varepsilon : W_0 \times W_0 \rightarrow \mathbb{R}$ and $b_0^\varepsilon : H_0 \times H_0 \rightarrow \mathbb{R}$ defined by

$$a_0^\varepsilon(\mathbf{v}, \mathbf{w}) := \int_0^\ell \mathbf{H}^\varepsilon \mathbf{a}(\mathbf{v}) \cdot \mathbf{a}(\mathbf{w}) dz^1, \quad b_0^\varepsilon(\mathbf{v}, \mathbf{w}) := (\varrho^\varepsilon A^\varepsilon \mathbf{v} \mid \mathbf{w})_{H_0},$$

correspond to the forms a_0 and b_0 . It can be shown by simple calculation that the problem (12), (13) is then equivalent to

$$\begin{aligned} \frac{d}{dt} b_0^\varepsilon(\partial_t \mathbf{u}, \mathbf{v}) + a_0^\varepsilon(\mathbf{u}, \mathbf{v}) &= (\overline{\mathbf{F}}^\varepsilon | \mathbf{v})_{H_0}, \quad \mathbf{v} = (\mathbf{v}, \psi) \in \mathcal{W}_0, \\ \mathbf{u}(0) &= \mathbf{u}_0, \quad \partial_t \mathbf{u}(0) = \mathbf{u}_1, \end{aligned}$$

where

$$\overline{\mathbf{F}}^\varepsilon = \int_{\varepsilon S} \mathbf{F}^\varepsilon dy^2 dy^3.$$

References

- [1] R. DAUTRAY AND J. L. LIONS, *Mathematical Analysis and Numerical Methods for Science and Technology, Volume 5, Evolution Problems I*, Springer-Verlag, Berlin, 1992.
- [2] R. JAMAL AND É. SANCHEZ-PALENCIA, *Théorie asymptotique des tiges courbes anisotropes*, C. R. Acad. Sci. Paris Sér. I Math., 322 (1996), pp. 1099–1106.
- [3] M. JURAK AND J. TAMBAČA, *Derivation and justification of a curved rod model*, Math. Models Methods Appl. Sci., vol. 9, no. 7 (1999), pp. 991–1014.
- [4] M. JURAK AND J. TAMBAČA, *Linear curved rod model. General curve*, in preparation.
- [5] XIAO LI-MING, *Asymptotic analysis of dynamic problems for linearly elastic shells – Justification of equations for dynamic membrane shells*, Asymptotic Anal., 17 (1998), pp. 121–134.
- [6] A. RAOULT, *Construction d'un modèle d'évolution de plaques avec terme d'inerte de rotation*, Ann. Mat. Pura Appl. (4), 139 (1985), pp. 361–400.
- [7] J. TAMBAČA, *One-dimensional models in theory of elasticity*, M.Sc. thesis, Department of Mathematics, University of Zagreb, 1999. (In Croatian).
- [8] J. TAMBAČA, *Evolution model of curved rods*, Ph.D. thesis, Department of Mathematics, University of Zagreb, 2000.